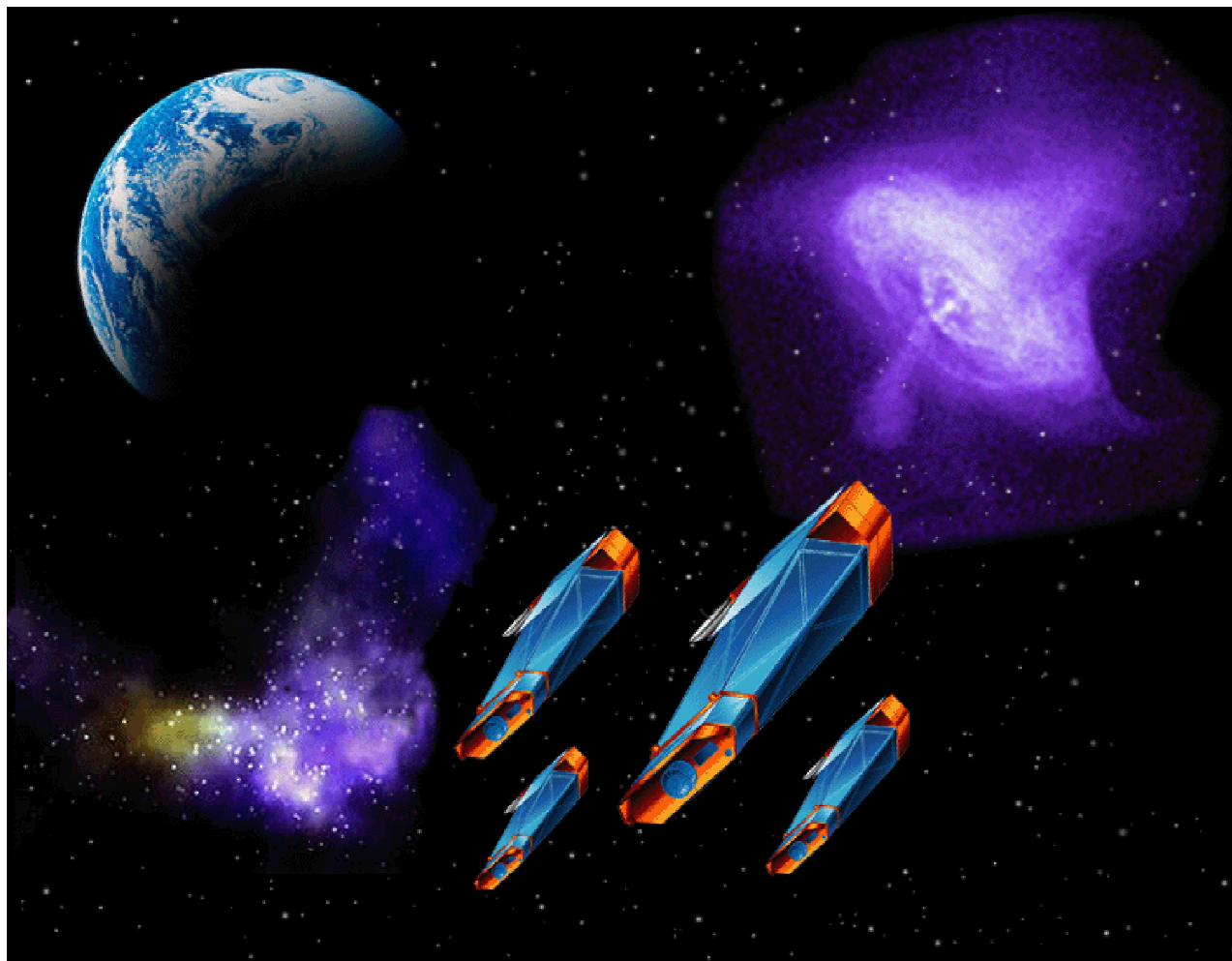




# Con-X Spectroscopy X-Ray Telescope

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Rob Petre, GSFC, SXT IPT Lead





# Constellation-X Requirements Flow Down

## Science Goals

Parameters of  
Supermassive  
Black Holes

Search for  
Dark Matter

Investigate  
Faint Sources

Plasma  
Diagnostics  
from Stars to  
Clusters

## Measurement Capabilities

### Effective area:

*15,000 cm<sup>2</sup> at 1 keV  
6,000 cm<sup>2</sup> at 6.4 keV  
1,500 cm<sup>2</sup> at 40 keV*

### Band pass:

0.25 to 40 keV

### Spectral resolving power (E/ΔE):

*≥ 300 from 0.25 to 6.0 keV  
≥ 3000 at 6 keV  
≥ 10 at 40 keV*

### System angular resolution and FOV:

*15 arc sec HPD and  
FOV > 2.5' (0.25 to 10 keV)*

1 arc min HPD and  
FOV > 8' (10 to 40 keV)

## Engineering Implications

### Effective area:

- *Light weight, highly nested, large diameter (1.6 m) optics*
- *Long focal length (8-10 m)*

### Band pass:

- 2 types of telescopes to cover energy range

### Spectral resolving power:

- *Dispersive and non-dispersive capability* to cover energy band

### System angular resolution and FOV:

- *Tight tolerances on telescope figure, surface finish, alignment*
- ≥ 30 x 30 array for x-ray calorimeter (pixels ~5")
- Cryocooler driven by array size and readout electronics

## Key Technologies

### High throughput optics:

- *High performance replicated segments and shells*
- *High reflectance coatings*
- *High strength/mass materials for optical surfaces*

### High energy band:

- Multilayer optics
- CdZnTe detectors

### High spectral resolution:

- 2 eV calorimeter arrays
- Coolers
- Lightweight gratings
- CCD arrays extending to 0.25 keV

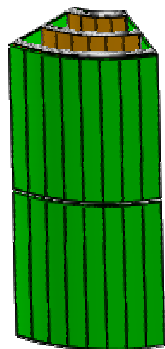
### Optical bench:

- Stable (time and temp.)
- High strength/low weight materials



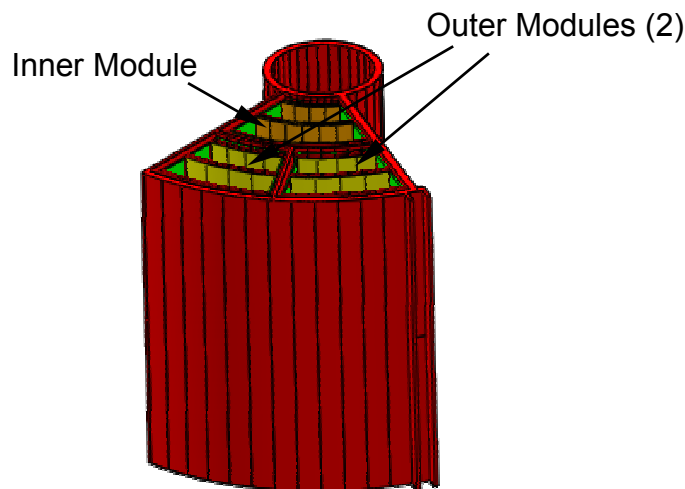
# SXT Mirror Phased Development

**Engineering Unit and  
Optical Alignment  
Pathfinder**



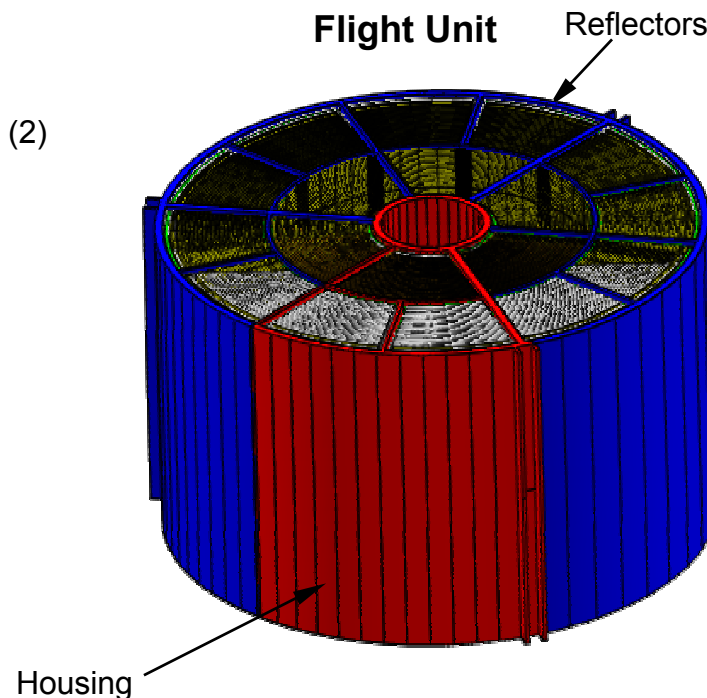
- Single inner module with
- 0.5 m dia. reflector pair (replicated from Zeiss precision mandrel)
  - Parabolic (P) and Hyperbolic (H) submodules
  - First modules to be aligned using etched silicon microcombs

**Prototype Unit**



- Flight scale assembly of
- 3 modules (2 outer and 1 inner)
  - Largest diameter same as for flight - 1.6 m
  - Each module has 3 to 9 reflector pairs
  - Demonstrates module to module alignment

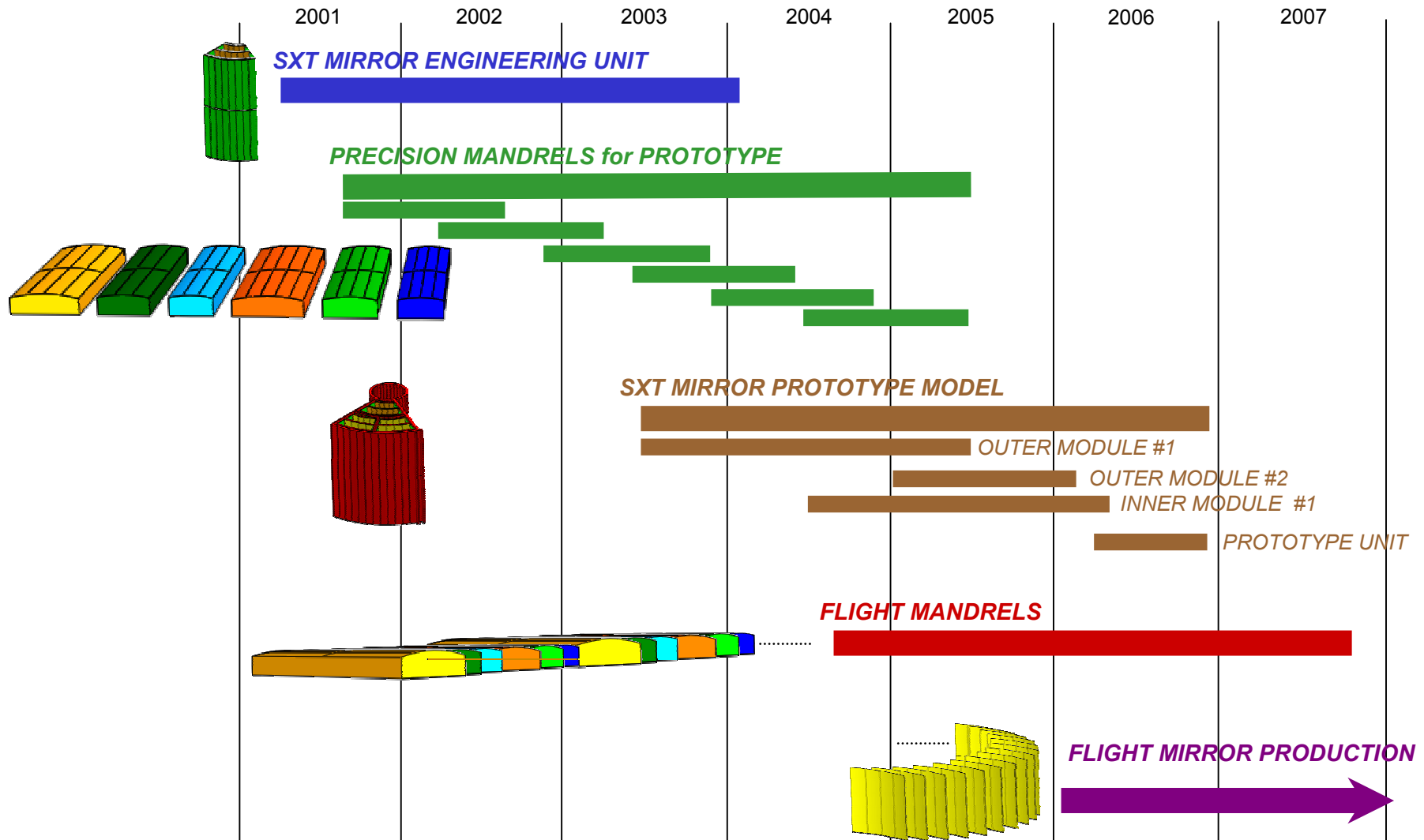
**Flight Unit**



- Full flight Assembly
- 1.6 m outer diameter
  - 18 Small Modules
  - 0.2 - 0.5 m segment length
  - 70 to 220 reflector diameters



# SXT Technology Roadmap





# SXT Prototype Development

	Optical Assembly Pathfinder		Eng. Unit	Prototype		
Configuration						
Module Type	Inner	Inner	Inner	Outer	Inner	Outer & Inner
Housing Material	Aluminum	Titanium	Composite	Composite	Composite	Composite
Focal Length	8.5m	8.5m	8.5m	10.0m	10.0m	10.0m
Reflector Length (P&H)	2 x 20 cm	2 x 20 cm	2 x 20 cm	2 x 50 cm (TBR)	2 x 50 cm (TBR)	2 x 50 cm (TBR)
Nominal Reflector Diameter(s)	50 cm	50 cm $\pm$	50 cm $\pm$	160 cm $\pm$ 120 cm $\pm$ 100 cm $\pm$	90 cm $\pm$ (TBR) 70 cm $\pm$ (TBR) 50 cm $\pm$ (TBR)	160 cm $\pm$ 40 cm $\pm$ 120 cm $\pm$ 70 cm $\pm$ 100 cm $\pm$ 50 cm $\pm$
Goals	<ul style="list-style-type: none"> <li>Align 1 optical surface pair (P&amp;H)</li> <li>Evaluate optic alignment techniques, optics assembly design &amp; process, &amp; optics metrology</li> </ul>	<ul style="list-style-type: none"> <li>Align up to 3 optical surface pairs (3P,3H)</li> <li>Evaluate tooling and alignment techniques for mass production schemes</li> </ul>	<ul style="list-style-type: none"> <li>Align 3 optical surface pairs to achieve &lt;10 arc sec.</li> <li>Environmental and X-ray test</li> </ul>	<ul style="list-style-type: none"> <li>Flight-like configuration outer module</li> <li>Largest optical surfaces</li> <li>Environmental and X-ray test</li> </ul>	<ul style="list-style-type: none"> <li>Flight-like configuration inner module</li> <li>Environmental (TBR) and X-ray test</li> </ul>	<ul style="list-style-type: none"> <li>Demonstrate module to module alignment</li> <li>Environmental and X-ray test</li> </ul>
Timeframe	Q4 of FY02	Q2 of FY03	Q1 of FY04	Q4 of FY05	Q3 of FY06	Q4 of FY06



# SXT Mirror Status

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- **Current effort focused on OAP1**
  - Glass forming and replication process well established
    - Figure of small reflectors (20 cm dia.) within a factor of 2 of requirement
  - Infrastructure nearly in place for 50 cm reflector forming and replication
    - Forming & replication mandrels delivered, oven & replication station operational, cutting fixture and spay booth being assembled
  - Al housings being assembled
  - Metrology & alignment approach defined; awaiting refurbished CDA from Bauer
- **Efforts underway to prepare for future OAPs and Prototype**
  - Large segment replication mandrels on order (Zeiss)
    - 1.6, 1.2, 1.0 m diameter; first delivery in August 2002
  - Prototype Si alignment structures developed by MIT
    - Existing structures compatible with OAP optical design
    - Provide inter-reflector alignment accuracy of  $\sim 0.1\mu\text{m}$
  - RFI released for large forming mandrels
  - Metrology facilities and other facilities being upgraded



Etched Si alignment  
microcomb



# SXT Mirror Development Team

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- **Multi-institutional participation**

- **GSFC**

- IPT leadership and management

- Mirror design and error budgeting

- Substrate forming and reflector replication

- Housing design and assembly

- Reflector and module metrology

- **MSFC**

- X-ray calibration

- Mandrel procurement support

- Mandrel metrology

- **MIT**

- Precision microlithography (alignment structures)

- **SAO**

- Management support

- Systems engineering

- **Industry partners**

- Zeiss (precision mandrels), Bauer (metrology), Rodriguez (forming mandrels)



# SXT Mirror Presentations

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Overview  
Technical and performance requirements  
Substrates  
Mounting and alignment  
Metrology  
Centroid Detector Assembly  
MSFC Activities

Rob Petre (GSFC)  
Bill Podgorsky (SAO)  
Will Zhang (GSFC)  
Jeff Stewart (GSFC)  
Dave Content (GSFC)  
Paul Glenn (Bauer)  
Steve O'Dell (MSFC)